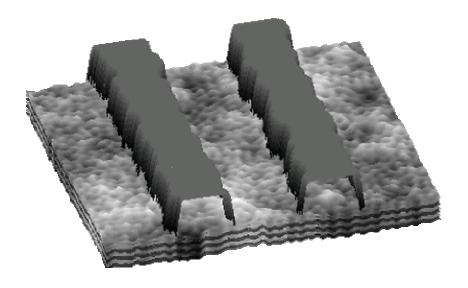


# How good will EUV masks need to be to meet LER requirements?

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#### **Outline**

**Problem** 

**Evidence** 

**Implications** 

Complications

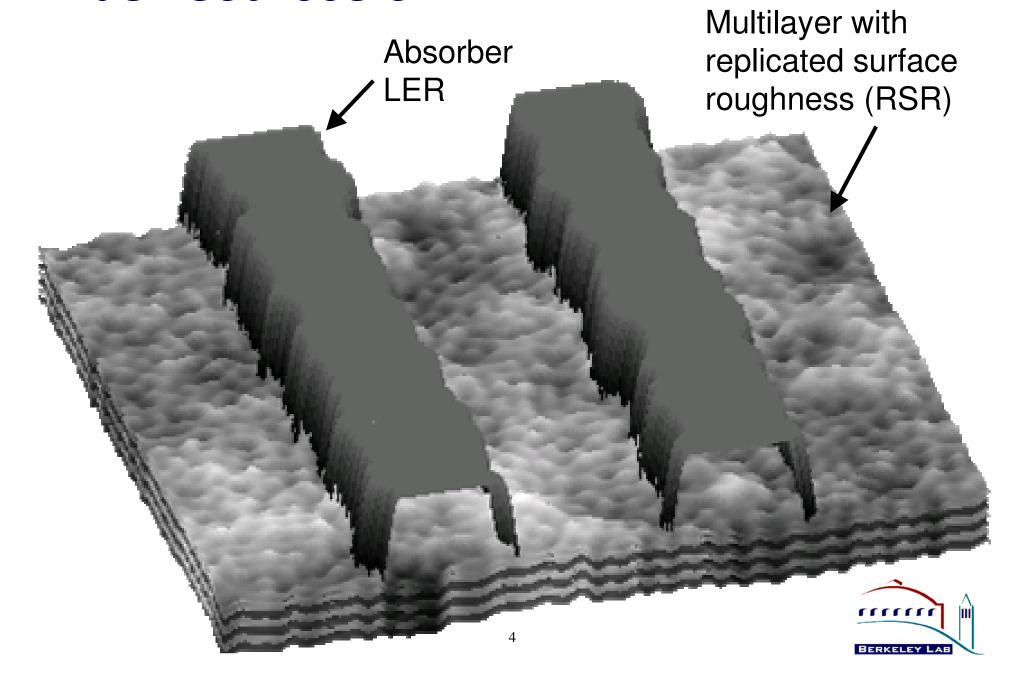
Summary



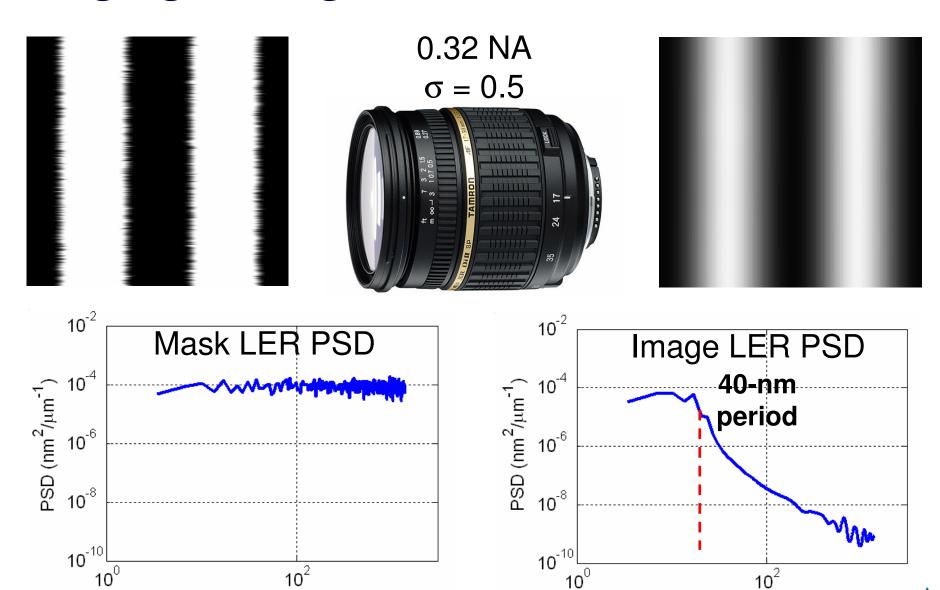
### The Problem



#### Mask sources of LER



#### Imaging demagnifies and filters mask LER



Appl. Opt. 42, 3390-3397 (2003)

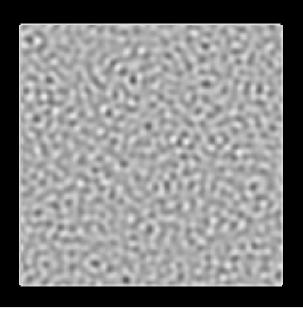
Spatial Frequency (lines per μm)

Spatial Frequency (lines per μm)

# Imaging transforms replicated surface (phase) roughness to intensity speckle



Contrast = 9%



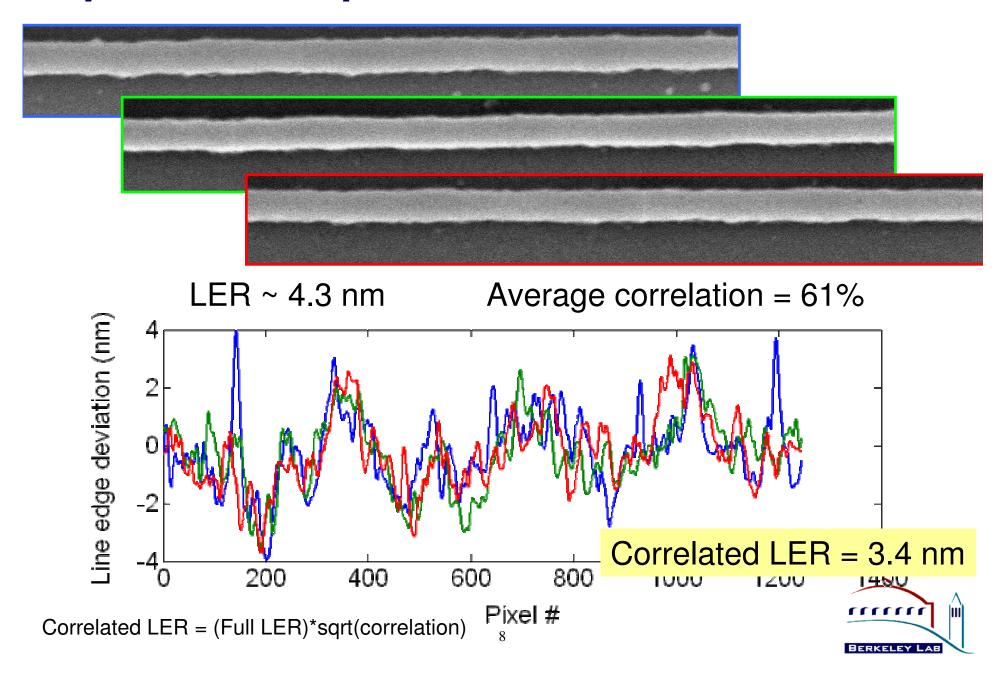
See Goldberg et al, Tuesday 12:20PM for experimental demonstration



# Experimental evidence



#### Exposure-to-exposure correlation observed



# Good agreement between measured correlated LER and modeled mask-induced LER

Configuration	Measured correlated LER (nm)	Modeled mask-induced LER (nm)
Mono, F=100 nm	$3.4 \pm 0.2$	3.0
Ann, F=100 nm	$2.7 \pm 0.3$	2.5
Ann, F=0 nm	$2.0 \pm 0.3$	1.4

<sup>\*</sup> Correlated LER = (Full LER)\*sqrt(correlation)
Uncertainty based on limited extent of correlation
measurement relative to bandwidth

9

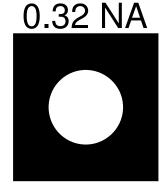


### *Implications*



### Modeling assumptions

**22-nm HP** 

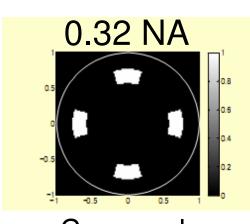


Disk  $\sigma = 0.5$ 

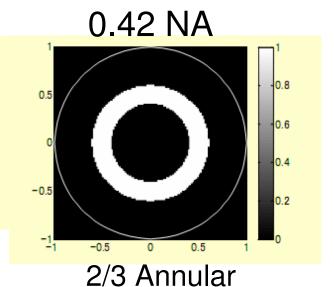
Ideal optic assumed in all cases

16-nm HP

From Canon's SPIE AL09 presentation



Cross-pole 
$$\sigma_{out} = 0.76$$
  $\sigma_{in} = 0.57$ 



 $\sigma_{\rm out} = 0.6$ 



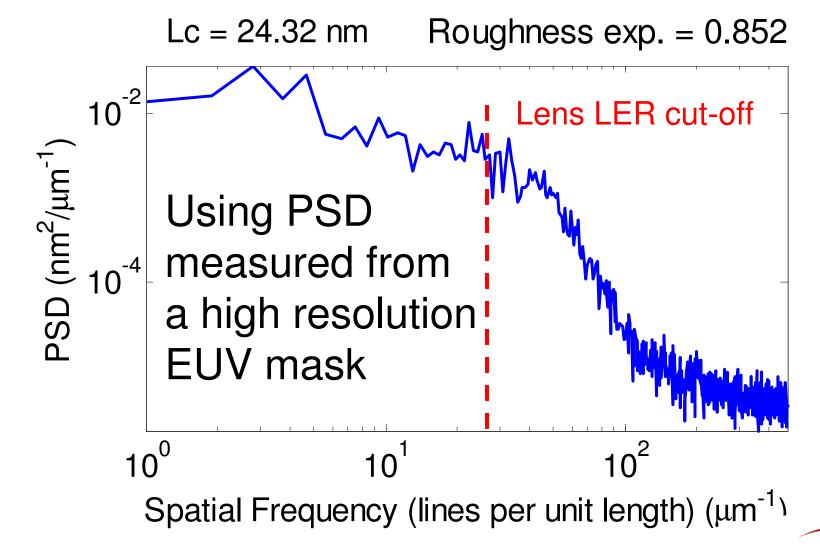
# Error budget allocation assumptions

Half pitch (nm)	22	16
Total image plane LWR (nm) <sup>1</sup>	1.8	1.3
Mask LWR contribution (nm) <sup>2</sup>	0.7	0.5
Allowable DOF reduction (%)3	30	30

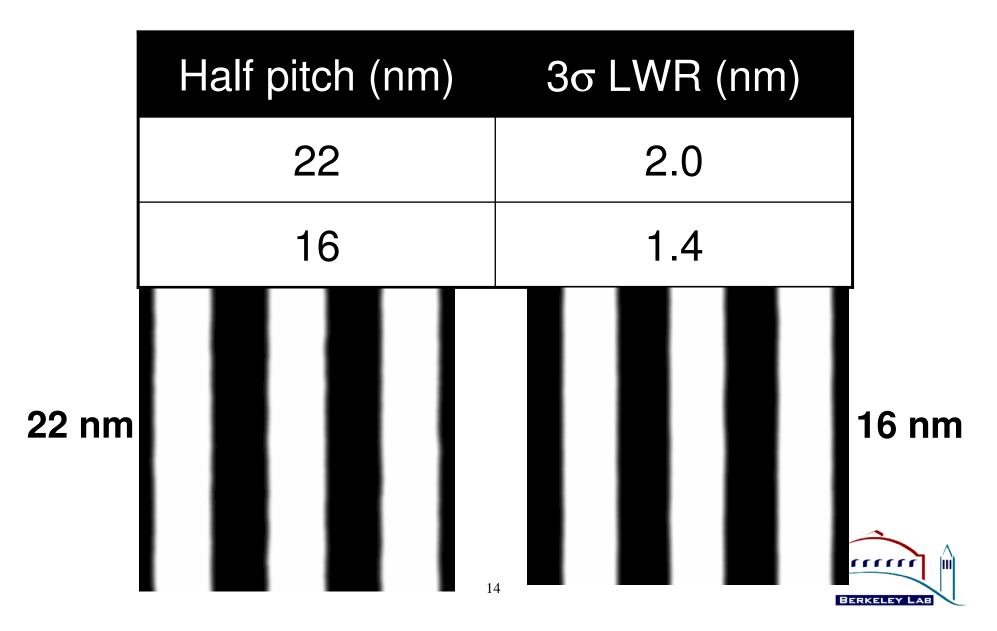
- <sup>1</sup> 8% of CD (from ITRS)
- <sup>2</sup> 10% contribution to total in quadrature
- <sup>3</sup> Reduction from the NILS = 1 DOF



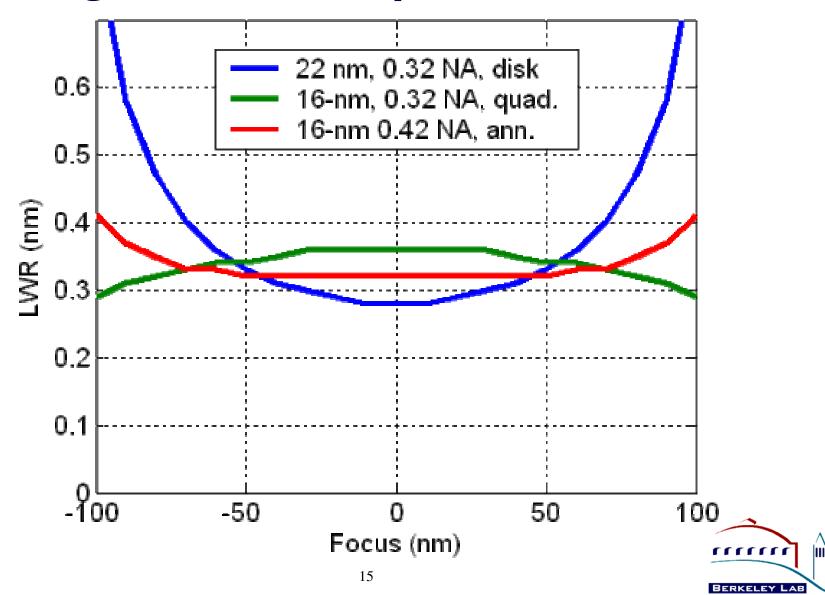
# Mask absorber LER coupling depends on mask LER PSD



## Mask LER magnitude based on 2008 ITRS



# Modeled image plane LWR resulting from ITRS spec mask LER

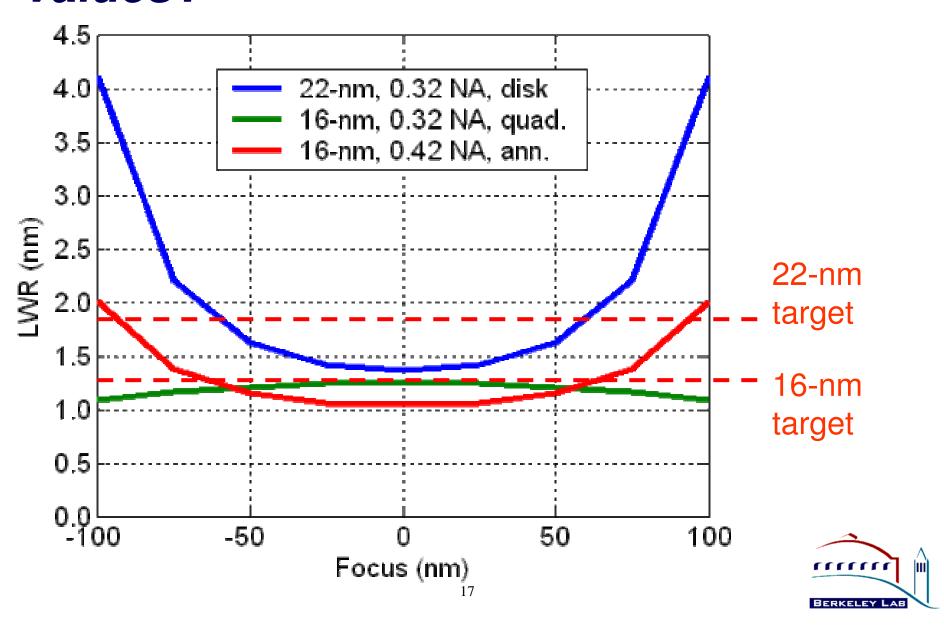


## What if we use expected mask LWR values?

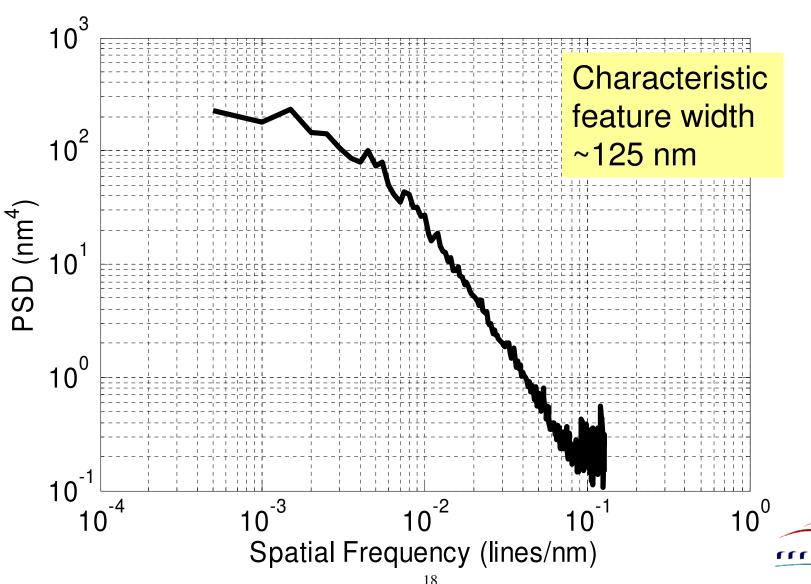
Half pitch (nm)	3σ LWR (nm)
22	8.0
16	6.0



## What if we use expected mask LWR values?



# Multilayer replicated roughness is generally low frequency



#### Roughness sensitivity @ 22-nm HP

#### **Litho Parameters**

- 0.32 NA
- Disk  $\sigma = 0.5$
- 22-nm half pitch
- Ideal optic

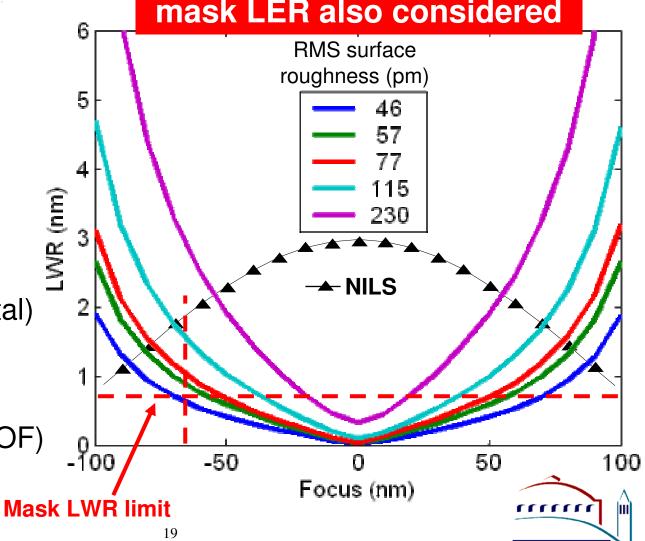
#### **LWR Limits**

- Total: 1.8 nm (8% of CD)
- Mask: 0.7 nm (10% impact on total)

#### **DOF** Requirement

• 130 nm (70% of NILS=1 DOF)

46-pm RMS surface roughness requirement if mask LER also considered



### Mask roughness limits summary

Configuration	RSR limit (pm)	RSR limit with mask LER (pm)
22-nm, 0.32 NA	46	46
16-nm, 0.32 NA	77	77
16-nm, 0.42 NA	77	57

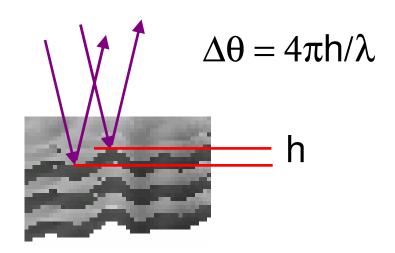


# Rough Capping Layer

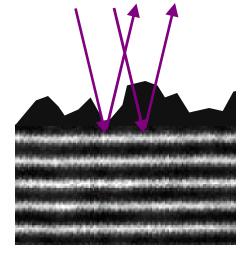


## With capping layer roughness, phase shift is no longer geometric, but refractive

#### RSR is geometric effect



Impact of capping layer roughness depends on capping material refractive index





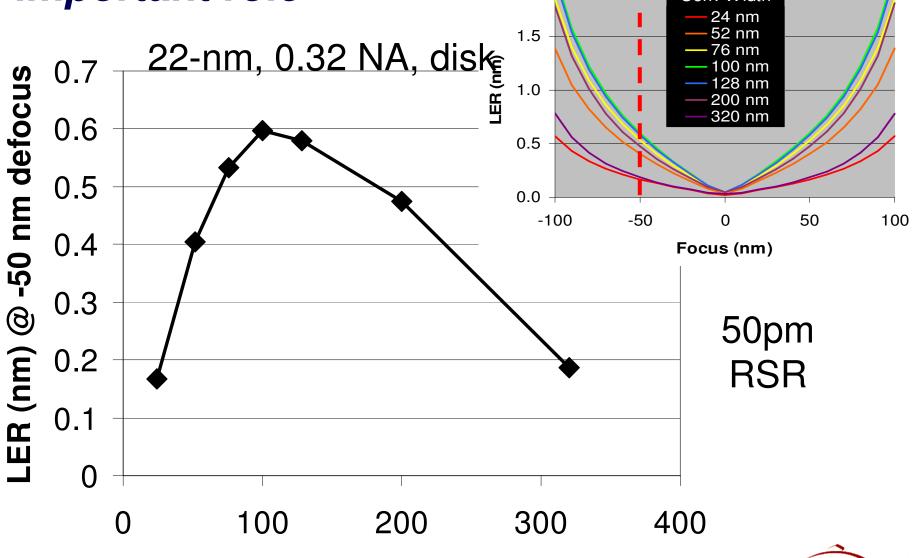
# Sensitivity to capping layer roughness highly dependent on material and much lower than RSR

Capping Material	Double Pass Phase Shift per nm of material	Roughness Equivalent to 50 pm RSR*
Si	0.002°	730 nm
Ru	6°	0.44 nm
С	2°	1.25 nm



Roughness correlation width plays important role

2.0 Corr. Width



Correlation width (nm)



#### Summary

- Replicated mask substrate roughness leads to image plane LER
- Current LER requirements indicate replicated roughness limits near 50 pm
- Predicted 50-pm RSR limit relies on achieving stringent absorber LER specs

### Acknowledgements

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